corrVandFlux

User Manual

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Prepared by:

3.1

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Document Change Log

Issue/ Rev.	Issue Date	Sections	Reason for Change
1	7 Oct 2020	All	Initial release
2	20 Sep 2023	All	Incorporates changes of software since issue 1
3	30 Sep 2023	All	Prefinal revisions
3.1	06 Nov 2023	5	Updated description of main control panel items
3.2	08 Jan 2024	5.1.13	Added output choice for photocurves

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1 Overview

The program corrVandFlux displays and analyzes data from the NASA Magnetospheric Multiscale (MMS) Mission related to spacecraft charging which have been preprocessed by the program mmsedpana in order to obtain spin averages of relevant quantities. In addition, MMS data may be correlated with solar activity data.

1.1 Instrument data

Inputs to the program include data from the instruments

Acronym	Name	Ref.	Parameter
ASPOC	Active Spacecraft Potential Control	[1]	Ion beam current
EDI	Electron Drift Instrument	[2]	Gun beam current
EDP	Electric Field Double Probes, consisting of:	[3]	
SDP	Spin Plane Double Probes	[4]	Spacecraft potential, Electric field
ADP	Axial Double Probes	[5]	Spacecraft potential, Electric field
FPI	Fast Plasma Instrument, consisting of:	[7]	
DES	Dual Electron Sensors		Electron moments, energy distributions
DIS	Dual Ion Sensors		lon moments, energy distributions

1.2 Input data base

The program mmsedpana has been used to generate spin average data of spacecraft potential, electric field, ASPOC current, electron and ion density, temperature, and current. The files in the distribution cover the time range 2015-09-01 to 2023-05-31. Due to the long processing time, the data in Fast Survey mode have been analysed in batches of 2 months, and for electron and ion data separately. The combination of electron and ion data has been performed by the program mmsedpoutput_merge_vel_ei_leo. The concatenation of the 2-month files has been performed by the program mmsedpoutput_merge_vel_ei_leo. The steps have been performed in the Unix environment at IWF (leo1). The results were copied to a Windows environment. Finally, the Fast Survey and Slow Survey data were merged by the program by the program mmsedpoutput_merge_vel_fastslow. The nomenclature of these files is mms*_out_*eandivfands.dat. The first lines of one of these files are reproduced below.

Time[UT] Phase12 E12ampl E34ampl Etotamp ExDSL EyDSL Vsc El.Dens El.Temp El.Curr ASP-EDI veGSEx veGSEy veGSEz lonDens lonTemp IonCurr viGSEx viGSEy viGSEz Time[UT] [deg] [mV/m] [mV/m] [mV/m] [mV/m] [V] [cm^-3] [eV] [uA] [uA] [km/s] [km/s] [km/s] [cm^-3] [eV] [uA] [km/s] [km/s] [km/s] 2015-09-01T09:52:06.527 59.617 -0.018 -0.261 0.262 -0.018 -0.261 3.003 0.650 259.67 2.845 19.712 -7.05 -21.41 0.763 5855.32 31.68 1.241 -0.67 2.58 8.45 0.609 272.42 2.728 19.710 2015-09-01T09:52:25.887 59.174 -0.159 -0.092 0.184 -0.159 -0.092 3.010 -0.71 -3.00 31.20 0.766 5712.50 1.230 0.20 0.92 6.47 2015-09-01T09:52:45.246 69.954 -0.109 0.171 0.202 -0.109 0.171 3.023 0.495 329.90 2.435 19.711 5.66 12.28 32.16 0.757 5698.14 1.213 3.56 5.81 5.98

The settings of mmsedpana can be seen in the control panels listed below.

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Tine No ----

X Selection of date and time for data and for tone calculation	-		×
Selection of date and time for data and for tone calculation SAVE CONTINUE WITHOUT SAVE			
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*repeated plots b Gaps betw, plots b			
TONE: use time All from year 2018 month 31 day 31 hr 30 min 30 sec 30 msec 300 to year 2018 month 31 day 31 hr 23 min 39 se	ac 100	NSEC	<u>}000</u>

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E12/34-bal No I E-phase shift No E-spikes Keep Ecomp-tone Keep Etot-to	ne Fr	°om comp.	-

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Selection of spacecraft potential source, offsets, corrections and de-tone parameters Smell Smell Smell CONTINUE WITHOUT SAME			
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Settings for density, current and potential reconstruction			
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Sample size 2048 Advance by sample fraction 4 Smoothing in freq. 1 in time 5			

1.3 Interaction between programs

Three programs working in the Windows environment analyse spin average data, define parameters for the reconstruction of plasma densities, and produce tables of the reconstructed values.

1.3.1 corrVandFlux

- Reads spin average data from a single spacecraft (mms*_out_edppeandiv.dat), orbit data, solar aspect angle data, and solar indices
- Calculates reconstruction parameters for a single magnetospheric region or for all regions together, and outputs files of these parameters. Available regions are: magnetosphere, magnetosheath, solar wind. Parameters are different for ASPOC OFF and ASPOC ON (nominal current). Therefore there are 6 parameter files in total.
- Produces plots with data from a single region or all regions together.
- Plots solar correlations.
- Calculates reconstructed densities.
- Uses a special temperature law for the combination of ASPOC ON, inside magnetosphere.
- Uses correction terms for the spacecraft potential in dependence of electric field, ion velocity, F10.7 cm flux, ion Mach number.
- Outputs parameter files to be used by corrVandFluxPredef.
- Outputs reconstructed density files for a single region and a single ASPOC state in a format compatible with outputs of multiple regions produced by corrVandFluxPredef.

1.3.2 corrVandFluxPredef

- This program serves to use the six parameter files produced by corrVandFlux to produce a single file with reconstructed densities covering all regions.
- Reads spin average data from a single spacecraft (mms*_out_edppeandiv.dat), orbit data, and solar indices
- No filtering of input data except for time
- Results near the region boundaries are interpolated in order to get smooth transitions
- Outputs reconstructed density files for all regions and ASPOC states together in a format compatible with outputs of single regions produced by corrVandFlux.

1.3.3 corrVandFluxRec

- Reads a reconstructed density file, orbit data, and solar data.
- Plots solar correlations.
- Plots correlations between measured and reconstructed data.

1.4 **Processing options**

Processing options include the following

- Plots or tables of data over time
- Plots of MMS data in the equatorial plane
- Correlation of MMS data with solar activity indices
- Correlation between spacecraft potential and plasma density
- Correlation between spacecraft potential and currents from plasma and ASPOC
- Correlation between spacecraft potential and electric field
- Correlation between electron and ion density
- Histograms of spacecraft potential and density
- Calculation of photocurve from dual spacecraft potential data in multiple function options
- Calculation of photocurve from spacecraft potential and plasma electron data in multiple function options
- Reconstruction of uncontrolled spacecraft potential based on the above fitting methods
- Reconstruction of plasma density based on the above fitting methods
- Reconstruction of plasma current based on the above fitting methods
- Plots of plasma current over spacecraft potential, plasma temperature, electron velocity, ion velocity, and ion Mach number
- Plots of particle density over spacecraft potential, electron velocity, and ion velocity
- Plots of reconstructed over measured particle density
- Plots of reconstructed over measured plasma current
- Plots of spacecraft potential over electron velocity, ion velocity, temperature, and ion Mach number
- Plots of electron velocity, ion velocity, and ion Mach number over temperature

1.5 Output

- Plots in PNG or Postscript format
- ASCII tables of input data together with reconstructed plasma data from the selected orbital regions.
- ASCII files of the parameters used for the reconstruction in the selected orbital region

2 System Requirements

Source codes are compatible with IDL 8.7 or higher. The program works best in a Windows environment.

3 Installation

3.1 Contents of distribution

The distribution comes as a zip file corrVandFlux.zip containing three directories.

doc	documentation including this document and the EDP data products guide
data	files with spin average data, orbit data, solar aspect angle of the spacecraft, and solar activity
source	IDL source code

3.2 Directories

Input data (spin average MMS data, orbit data, solar aspect (tilt) angle of the spacecraft, and solar activity) should be located in the same directory.

If the option to generate an output file with reconstructed density data is selected, the user will be prompted for the output directory and file name.

3.3 Step-by-step installation procedure

- Unpack the zip file
- Move the files in the directory "source" to the appropriate path of the IDL distribution for subsequent compilation and execution.
- Move the files in the directory "data" to the final location.

Compile the sources and execute them, or load the sources into the IDL development environment and execute them.

3.4 **Program usage**

• At the beginning the following window entitled "Select the directory with MMS and solar data" will appear for the input of a directory which contains all input data:



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Thereafter the user shall select the primary MMS data file for analysis. The following window entitled "Select the file with MMS data" will appear. This file is mandatory. Three versions of spin average data files can be selected: Files with Fast Survey mode data only, files with Slow Survey mode data only, and files with merged Fast and Slow Survey mode data. The files follow the nomenclature mms*_out_<fastslow>edppeandiv<mergeID>.dat. "eandi" stands for the fact that the contain both electron and ion data. Examples: mms1_out_fastedppeandiv.dat : MMS1 Fast Survey data mms2_out_slowedppeaniv.dat: MMS3 Fast and Slow Survey merged data

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늘 Desktop	10	B mms1_out_eandi.dat	19.04.2020 17:10	DAT-Datei	49
Dokumente		@ mms1_out_eandi_filt.dat	16.06.2020 10:13	DAT-Datei	49
Downloads		mms1_out_eandi_filt_recoff.dat	20.09.2020 21:19	DAT-Datei	18
Musik		mms1_out_eandi_filt_recon.dat	16.09.2020 18:08	DAT-Datei	17
The Midage		@mms1_out_eandi_filt_reconoff.dat	20.09.2020 21:27	DAT-Datei	36
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Uindows (C:)		mms1_out_eandi2019.dat	06.06.2020 00:07	DAT-Datei	2
🗙 mms (\\leonas2) (M:)		mms1_out_eandikurz.dat	20.04.2020 10:56	DAT-Datei	
WF-Shares (N:)		@ mms1_out_mean.dat	22.01.2020 10:43	DAT-Datei	15
≈x Home (O:)	~	mms1 out mean dce dat <	. 22 01 2020 10:47	DAT-Datei	35
Dateiname:	-		~ mms*c	out*.dat	<i></i>

 Thereafter the following window entitled "Select a secondary file with MMS data or none" will appear:

Select a secondary file with MMS of	data,	or none				×
← → × ↑ 📜 + Data >	MM	S > ASPOC > cost > ASAP15 > data >	v U	"data" du	irchsuchen	م
Organisieren • Neuer Ordr	er					. 0
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Datei <u>n</u> ame:			~	mms*ou Öffn	t*.dat en Abł	↓ prechen

Here the user may select a MMS data file of a different spacecraft in order to perform comparisons between controlled and uncontrolled potentials. If such a comparison is not requested, pressing "Cancel" ("Abbrechen") leads immediately to the next window.

• Thereafter the user may select a file with partial moments data for additional analyses in the following window entitled "Select a file with partial moments data or none". These files are optionally produced by mmsedpana when partial moments analysis is enabled.

elect a file with partial moments	data,	or none			×
() 🕥 🛧 📜 Data :	MMS	5 > ASPOC > cost > ASAP15 > data >	✓ ບ °data	durchsuchen	م ر
Organisieren • Neuer Orc	Iner			Ⅲ •	. 0
S Dieser PC	^	Name	Datum	Тур	Größe ^
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besktop	10	mms1_out_eandi.dat	19.04.2020 17:10	DAT-Datei	49
B Dokumente		mms1_out_eandi_filt.dat	16.06.2020 10:13	DAT-Datei	45
Downloads		mms1_out_eandi_filt_recoff.dat	20.09.2020 21:15	DAT-Datei	18
Musik		mms1_out_eandi_filt_recon.dat	16.09.2020 18:0	DAT-Datei	17
Videos		mms1_out_eandi_filt_reconoff.dat	20.09.2020 21:2	DAT-Datei	36
Videos		mms1_out_eandi_reduced.dat	19.04.2020 17:10	DAT-Datei	47
Windows (C:)		mms1_out_eandi2019.dat	06.06.2020 00:0	DAT-Datei	2
🐋 mms (\\leonas2) (M:)		mms1_out_eandikurz.dat	20.04.2020 10:50	DAT-Datei	
碱 IWF-Shares (N:)		@mms1_out_mean.dat	22.01.2020 10:43	DAT-Datei	15
Home (O:)	~		22 01 2020 10-4	DAT-Datei	35 *
Dateiname			~ mms	'out*.dat	<i></i>
outeDanie	·		111115	out lost	

If a partial moments file is selected, all analyses involving FPI plasma data will be based on partial moments with a selectable lower energy limit.

• Thereafter the main control panel shown below will appear. For further instructions refer to section 5.

E Parameter Selection ×
EXECUTE WITH CURRENT DATA LOAD NEW DATA REFRESH PANEL END PROGRAM Input: mms1_out_edppeandiv_2019.dat Primary SIC: mms1_Secondary SIC: none
Output Plots to screen only · Process density data · Solar wind little Any · Dump data No · Hide His No · Plot style Std · OR size Fixed · Char size 15
Start time: year 2015 month 11 day 01 hour 00 minute 00 second 00 End time: year 2023 month 06 day 01 hour 00 minute 00 second 00
Solar Data F10.7 V Analyse Electrons v above eV 6.52*all v Plot Solar correl No v Vscvs E No v Time series No v Corr. E-I No v InThy No v Vsc-Tv No v Detrend Yes v measured corr with No v
Solar Activity. Irradiance>0.0000 <0.0008 F10.7 Fkue>60.0 <300.0 Sunspots>0.0 <500 Tit(deg)>0.00 <500
Orbit Re 10.00 < 3000 Phase(Sun=0)= 00 < 3600 In&Out > Orbit map None > Scale Limits > Resol. 0.5Re > Central Msheath@X&+Y&-YGSE 120 200 100 Region Al > Width@Y=08Y=20[25] 70
Color label is Temperature v Colour is occurrence No v Label time plots No v #labels in n-V-plots 0 in V20 40 and n0080 1000 Scale ranges Limits v 2nd param f. solar corret. None v
Correlate Vsc with Density & current 🗸 As j Use (tb) 🗸 As v Use Total 🗸 with v times 1.00 ASPOC step analysis: time window(5)(60.0 I +window)7.0 Max gap(4)
n &I-T-exponent: 0.25 applies for T< 50.0 and exp. 0.00 for higher T up to: 0.0 Enhance n&I at T< 50.0 by 0.00 Min T for I-catc 0.0 FPI spectra No v Max energy step 25 Regres w/V No v
Manual correction of Isoc No v gr points 0 Vsc 200 3.00 5.00 10.00 20.00 40.00 Factors 100 1.00 1.00 1.00 1.00 1.00
Histograms of SC-Potential vs. None v #bins. 10 #param steps. 14 Scatterplots 1 x points. 50 Symbol size 0.3 #bins. 300 cutoff 30 #equal weight bins. 1 Omit data near NaN Yes v
Correc. for Vsc. E-field[tm] 0.00 Sun 0.0000 v [mV skm] 0.0000 Mach 0.000 Multiply V-corr by V No v [g[T] 0 0.0000 [g[T] n 0.0000 Tithresh 300.000 Fudge factor for maxcurrent 1.00 for [plasma 1.00
Show VV power fit in IV-piol No 🗸 coefs jph 65.00 bph 1 2000 Use Vsc dev. from smoothing x spins (0-no) x 0 Smooth dVidE (0-no) 501 [dVdE]scale+ 1.0 Divide dVdE by V No 🗸 Output photocurve No 🗸
Fit I or n vs V on set polynomial No v #coels 5 coels 19.72 -18.72 240 3.53 -1.11 0.00 0.00 Plot poly Yes v Set No v ph031.90 V0 or exp1610 Set No v ph10.00 V11000 ph2000 V2 1.000
Fit I vs V No Maxw fit Mpfit at Terms 1 Limits Yes Method Simple Error exp 0 in Y Break V Var 60 140 T for n-fit Var at 2v/200 7v/400 10.5v/700 20v/300.0 ftx 150.0 SETTINGS FOR SPACEGRAFT 1
Efields-0.0 < 20.0 Vsc-2.0 < 50.0 Vsc-scale Logarithmic v Iaspoor-1.0 < 1.0 Density (no mi< 999.000 AND (no mi< 5.0 torn tar. 5.
Electrons New=0.0880000 <[1000.000 Tex=5.0 <[1000.00 tex=0.100000 <[1000.0 we=0.0 <[1000.0 we=0.0 <[1000.0 tors.Nex=0.010000 <[1000.00 Tex=5.0 <[1000.0 tors.Nex=0.010000 <[1000.0 tors.Nex=0.010000 Tex=5.0 tors.Nex=0.010000 <[1000.0 tors.Nex=0.010000 Tex=5.0 tors.Nex=0.010000 <[1000.0 tors.Nex=0.010000 Tex=5.0 tors.Nex=0.01000 Tex=5.0 tors.Nex=0.010000 Tex=5.0 tors.Nex=0.01000 ex=5.0 tors.Nex=0.01000 e
SETTINGS FOR SPACECRAFT 2 OR FOR UNCONTROLLED DATA IF APPLICABLE
Plot Vsc2(unc) over Vsc1(cont) OR Vsc1 at ASPOC changes No 🗸 Iaspoc2>[10] <[10] (Vsc2 only) Vsc2>[20] <[50.0] Vsc2 scale Logarithmic 🗸 Vsc2 correction[0.00]
Fit None Maxw break V Var 60 Error exp. 0 v Nat coefs 70.00 1.80 5.00

4 Output Data

4.1 Plots

By default the program outputs plots to the display. By selecting "Plots to XXX files" in the drop list for output files the user may generate plot files containing the same graphics as on the display. By default, the program proposes a file name consisting of the name of the main input file appended by an underscore and a code given in section 4.1.1. The format of the plot files is PNG or Postscript. The user may change the file name.

Code	Content
corr	spacecraft potential versus solar index
corri	plasma current versus solar index
corrn	spacecraft potential versus solar index, normalized
curr	plasma current over time
dens	particle density over time
djde	dj/dE vs E
dvde	d(Vsc)/d(E) over time
efld	total electric field over time
etvsca	spacecraft potential versus total electric field
evsc	total electric field versus spacecraft potential
f107	Penticton radio flux over time
fivsc	modified plasma current versus spacecraft potential
iasp	reconstructed ASPOC current versus spacecraft potential
imrec	reconstructed over measured plasma current, in single data point format
imrecb	reconstructed over measured plasma current, in histogram format
ioverm	plasma current over ion Mach number
iovert	plasma current over plasma temperature
ioverve	plasma current over electron velocity
iovervi	plasma current over ion velocity
ivsc	plasma current versus spacecraft potential, in single data point format
ivscb	plasma current versus spacecraft potential, in histogram format
movert	ion Mach number over temperature
neni	electron versus ion density
nhist	histogram of plasma density
nmrec	reconstructed over measured particle density, in single data point format
nmrecb	reconstructed over measured particle density, in histogram format
novert	particle density over temperature
noverve	particle density over electron velocity
novervi	particle density over ion velocity
nvsc	particle density versus spacecraft potential, in single data point format
nvscb	particle density versus spacecraft potential, in histogram format
pnrvsc	partial particle density versus residual spacecraft potential
pnvsc	partial particle density versus spacecraft potential
pos	orbit map
regr	regression coefficient of potential with partial particle density
see	solar irradiance over time
seef	solar flux index, filtered
spect	particle spectrogram
spot	American sunspot number over time

4.1.1 Plot file types and file names

Code	Content
SW	solar wind over time
tvsc	temperature versus spacecraft potential
veovert	electron velocity over temperature
vhist	histogram of spacecraft potential
viovert	ion velocity over temperature
VSC	spacecraft potential over time
vscoverm	spacecraft potential over ion Mach number
vscovert	spacecraft potential over temperature
vscoverve	spacecraft potential over electron velocity
vscovervi	spacecraft potential over ion velocity
vuvc	uncontrolled versus controlled spacecraft potential

4.2 Tables

By selecting "Plots to screen and data to file" or "Plots to screen and all data with filtered ASPOC to file" in the drop list for output files the user may generate an ASCII file containing time series of:

spacecraft potential (input)

total electric field (input)

electron density (input)

electron temperature (input)

electron current (input)

ASPOC current minus EDI current (input)

sum of electron and ASPOC current as result of the fitting

density reconstructed from spacecraft potential and measured temperature

density reconstructed from spacecraft potential and assumed temperature of 10 eV

density reconstructed from spacecraft potential and assumed temperature of 100 eV density reconstructed from spacecraft potential and assumed temperature of 1000 eV ion velocity (input)

ion Mach number (input)

region ID (0=any, 1=magnetosphere, 2=solar wind, 3=magnetosheath)

ID for parameter set used

ID for interpolation between regions (0=no, 1=yes)

ID showing that interpolated values are replaced by values derived from measured densities

In the distribution these output files have names in the nomenclature mms<N>_rec_*.dat. Depending on the selection of orbit regions, these files contain data from within these regions only. Such files can be used later to plot reconstructed parameters from the given orbit region by the program corrVandFluxRec. In addition, a parameter file (see 4.3) can be generated Some lines of one of these files are reproduced below.

Source: mms1_out_edppeandiv. Data include velocities Analysing electrons Limits (from/to): 2015-11-01T00:00:00/2016-11-01T00:00:00 E-field: 0.00/10.00 S/C potential: 2.0/50.0 EI. Density: 0.080/1000.000 EI. Temp.: 5.0/1000.0 EI. Current: 0.100/1000.000 ASPOC current: -1.00/1.00 LyA: 0.0000/0.0098 Orbit radius: 10.00/30.00 Orbit phase (Sun=0): 0/360 | ne-ni | <999.0 OR ne/ni<2.00 In Magnetosphere Transition width: 2.5 T-exponent: -0.25 applies for T< 50.0 and exp: 0.00 for higher T up to: 0.0 E-field correction term for Vsc: 0.00 v(ion) correction term for Vsc: 0.00 Fudge factor for maxcurrent: 1.00 Fudge factor for Iplasma: 1.00 Use set low-E-Maxw for Maxw-fit: No coefs:j0: 31.90 VO: 1.610 Fit I vs V: Power i~V^x or n vs V: No #Terms: I Limits: Yes Method: Simple Error exp: 0 in: Y Break V: Variable Vbreak le: 6.0 Vbreak2e: 14.0 T for n-fit: Variable at 2V: 20.0 at 7V: 40.0 at 10.5V: 70.0 at 20V: 300.0 fixed: 150.0 jph=30.762*V^(-1.708) T = 10.000 : n=213.832*V^(-1.708) T = 100.000 : n=67.620*V^(-1.708) T=1000.000 : n=21.383*V^(-1.708) END OF PARAMETERS Time[UT] Vsc Etotamp El.Dens El.Temp El.Curr ASP-EDI le+afit neTmeas neTOOIO neTOIOO neTIOOO Vion Machlon IDs [V] [mV/m] [cm^-3] [eV] [uA] [uA] [uA] [cm^-3] [cm^-3] [cm^-3] [km/s] [I] R P I F Time[UT] 2015-10-12T10:30:41.991 10.908 3.487 31.264 50.45 63.113 0.000 28.249 14.818 33.283 10.525 3.328 190.01 0.81 1 0 0 0 2015-10-12T10:31:01.335 11.197 3.849 31.357 49.32 62.662 0.000 28.330 15.030 33.378 10.555 3.338 188.12 0.87 1 0 0 0 2015-10-12T10:31:20.679 9.671 4.261 31.491 51.35 64.035 0.000 28.776 14.961 33.904 10.721 3.390 187.90 0.82 | 0 0 0

4.3 Parameter data files

Different orbital regions and ASPOC states require different sets of parameters for density reconstruction. Parameters for a single region and ASPOC state are calculated by corrVandFlux. These parameters are written into small data files, which are then used to produce a combined file of reconstructed data valid for all regions and ASPOC states. The header of the data file reproduced in the previous section represents an example for such a parameter data file. For the program corrVandFluxPredef to work properly, corrVandFlux must produce exactly 6 parameter files in the nomenclature:

mms*_para_<ASPOCstate>_<regionID>_<startdate>_<enddate>.txt with ASPOCstate = "on_" or "off"

ASPOCstate = "on_" or "off"

regionID = "msphere", "solarwd", or "msheath"

An example of such a parameter file is reproduced below.

I ; spacecraft number 2015 ; start year II; start month l ; start daty 0; start hour 2021 ; end year II; end month I; end daty 0; end hour 19.00 ; ASPOC min current 21.00 ; ASPOC max current 10.00 ; min orbit radius 30.00 ; max orbit radius 0.00 ; min orbit phase 360.00 ; max orbit phase 0; anyinoutbound I ; region selector 2.50 ; region transition width at Y=0 7.00 ; region transition width at Y=20 12.50 ; xgsedist (parabola of central magnetosheath) 20.00 ; pygsedist (parabola of central magnetosheath for +Y) 18.00 ; nygsedist (parabola of central magnetosheath for -Y) 0 ; as j use th | th+ve | th+vi | SQRT(th^2+ve^2) | SQRT(th^2+vi^2) 0 ; as v use Total | SpinPlane | Axial | X-GSE | Y-GSE | Z-GSE 1.00; with v times 0.000 ; E-field correction term for Vsc 0.0000 ; ion velocity correction term for Vsc 0.0000 ; Sun correction term for Vsc 0.0000 ; ion Mach number correction term for Vsc 0 ; solar parameter ID 5.90 ; sunlit area 10.03 ; surface area 2; powerrmsi I; ntermsrms 2; vsccorrwith 10.00 ; assumed temperature 1 100.00 ; assumed temperature 2 1000.00 ; assumed temperature 3 -0.25 ; n&I-T-exponent 50.00 ; applies for T< 0.00 ; and exp: 0.00 ; for higher T up to: 50.00 ; enhance n&l at T< 0.00 ; by 0.00 ; Min T for I-calc 0.00 ; lg(T)lo 0.00 ; lg(T)hi 300.00 ; Tthresh 0; Use set low E-Maxw ID 31.90 ; coefs:jph0 1.610 ; VO I ; T for n-fit ID 20.00 ; at 2V 40.00 ; 7V

70.00	; 10.5V
300.00	; 20V
150.00	; fix:
0	; method ID
1.00	; fudge factor for maxcurrent
1.00	; fudge factor for Iplasma
	; Currents
Power	; law
46.936	; jphO
-1.873	; exponent of V
	; Densities for fixed temperatures
10.00	; Temperature for the following parameters
Power	; law
55.300	; n0
-1.873	; exponent of V
100.00	; Temperature for the following parameters
Power	; law
17.487	; n0
-1.873	; exponent of V
1000.00	; Temperature for the following parameters
Power	; law
5.530	; n0
-1.873	; exponent of V

5 Control Panel Items

The program checks for most possible inconsistencies between entries in the control panel. However, bad entries are not always checked and the program might crash in these cases, for example if letters are entered into a field that requires a number. Possible inputs in the main control panel are numerical values, selections in drop lists, and buttons.

5.1 Main Group

 5.1.1
 Line 1

 Type:
 Button

 Label:
 EXECUTE WITH CURRENT DATA

 Description:
 Runs the program with the current selection of input files.

Type:ButtonLabel:LOAD NEW DATADescription:Allows the user to select new input files, but by default from the same directory
as at the start of the program. All other values in the control panel are kept.

Type: Button

Label: REFRESH PANEL

Description: Refreshes the main control panel. Thereby the energy ranges in the drop list for partial moments and the spacecraft names are updated.

Type:	Button
Label:	END PROGRAM
Description:	This button ends the program execution.

Labels at the right of the buttons indicate the primary input file and the selected spacecraft.

5.1.2 Line 2

Drop list Type: Label: Output Variable: makeoutput Defines output options for tables and plots Description: Selection: Plots to screen only Plots to screen only, no output in files Plots to screen and data to file Plots to screen, output of tabulated data including reconstructed densities (see section 4.2). The output file contains the filtered time and orbit intervals only. Plots to screen and all data with filtered ASPOC to file Plots to screen, output of tabulated data including reconstructed densities (see section 4.2). The output file contains all time stamps of the input file for which filtered ASPOC status is valid. Plots to screen and PNG files Plots to screen, and output of plot files in PNG format Plots to PNG files No plots to screen, output of plot files in PNG format Plots to PS files No plots to screen, output of plot files in Postscript format.

Type: Label: Variable: Description: Selection: all data density dat	Drop list Process withfpi Defines whether all input data input are processed or only those records which contain valid FPI data. All records in the input file are processed a Input records without valid FPI data are ignored
Type: Label: Variable: Description:	Drop list Solar wind filter swid This entry uses a beta version of a filter for solar wind conditions. It uses the conditions in the data file "mms1_20150901_0000_0000_dis_8_24_6_all.dat" which contains a status variable with 4 possibilities: inside solar wind, outside solar wind, undetermined, no data available. By default the user should enter "Any". Note that this data file has not been updated and should not be used when analysing data after 30 August 2021.
Selection:	
Definitely c	utside Filter for times being definitely outside seler wind
Outside or	no data
	Filter for times without available data or being outside solar wind
Definitely in	nside
	Filter for times being definitely inside solar wind

Inside or no data

Filter for times without available data or being inside solar wind

Any

Solar wind filter is not applied

Type: Label: Variable: Description: Selection:	Drop list Dump data dodump Allows to print test data in the IDL development environment.
No Yes	No output of test data Output of test data
Type: Label: Variable: Description: Selection:	Drop list Hide fits suppress Allows to print test data in the IDL development environment.

No Fitted lines are shown in all plots Yes Only data points, but no fitted lines are shown in the plots

Type: Label:	Drop list Plot style
Variable:	plotstyle
Description:	Defines the style of plots.
	Standard plot style, fixed plot size. By choosing a larger character size than
	1.5. the text is larger but may spill outside the plot window.
	Standard plot style, variable plot size: By choosing a larger character size than 1.5, the plot size is automatically increased if the character size exceeds 1.5. Publication plot style: character size is set to 2.5 and the plot size is increased accordingly
Selection:	
Std	Plots are produced in their standard format for working purposes
Pub	Plots are produced in a format suitable for publication
Type:	Drop list
Variable [.]	dovariableplotsize
Description: Selection:	Defines the size of plot windows, typically 800*600 pixels.

Fixed	Plots are produced in fixed size, without taking into account the size of the
	characters.
Variable	Plots are produced in variable size depending on the size of the characters.

Type:	Number
Label:	Char size
Variable:	cs
Description:	Defines the size of the characters in the plots. The standard value is 1.5.

5.1.3 Line 3

This line the user shall input the time range of the analyses. Note that the time range of the input data provided in the distribution is from 2015-09-01 to 2023-05-31.

Type:	Number
Label:	Start time: year
Variable:	yrmin
Description:	Year of the start of the time interval.

Туре:	Number
Label:	month
Variable:	momin
Description:	Month of the start of the time interval (1 12).

Type:	Number
Label:	day
Variable:	damin
Description:	Day of the start of the time interval (1 31).

Type:	Number
Label:	hour
Variable:	hrmin
Description:	Hour of the start of the time interval (0 23).
Type:	Number
Label:	minute
Variable:	mimin
Description:	Minute of the start of the time interval (0 59).
Type:	Number
Label:	second
Variable:	semin
Description:	Second of the start of the time interval (0 59).
Type:	Number
Label:	End time : year
Variable:	yrmax
Description:	Year of the end of the time interval.
Type:	Number
Label:	month
Variable:	momax
Description:	Month of the end of the time interval (1 12).
Type:	Number
Label:	day
Variable:	damax
Description:	Day of the end of the time interval (1 31).
Type:	Number
Label:	hour
Variable:	hrmax
Description:	Hour of the end of the time interval (0 23).
Type:	Number
Label:	minute
Variable:	mimax
Description:	Minute of the end of the time interval (0 59).
Type:	Number
Label:	second
Variable:	semax
Description:	Second of the end of the time interval (0 59).

5.1.4 Line 4

Type: Label: Variable:	Drop list Solar Data sunselect
Description:	Defines the solar activity parameter for correlations. Note that solar activity data are available in the period 2015-09-01 to 2023-06-30.
Selection:	
SEE	TIMED SEE L3A data of solar irradiance from https://cdaweb.gsfc.nasa.gov/cgi- bin/eval2.cgi?dataset=TIMED L3A SEE&index=sp phys
F10.7	Penticton Solar Radio Flux at 10.7 cm from https://lasp.colorado.edu/lisird/data/penticton_radio_flux/
Sunspots	American Relative Sunspot Number from https://lasp.colorado.edu/lisird/data/american_relative_sunspot_number_daily/
Type:	Drop list

Drop list
Analyse
species
Selects the particle species to be analysed.
Electron moments measured by FPI/DES
Ion moments measured by FPI/DIS

Type: Label: Variable: Description:	Drop list above eV fpimomsfrom Determines the energy level in FPI data for partial moments calculations. The selected energy is the lower boundary for the integration of flux over energy. This entry is only valid if partial moments data are selected. Note that the energy levels are default ones. The actual levels in the selected data may differ. This selection is only valid if a file with partial moments has been read, see section 3.4.
Selection: all energy valu	All energy levels are used
	Selected lowest bin of the energy
Tuno	Drop list

Type:	Drop list
Label:	Plot: Solar correl.
Variable:	plotsun
Description:	Allows to plot correlations between various quantities and the selected solar index.
Selection:	
No	No plot
Yes	Plots correlations between various quantities and the selected solar index

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Type: Label: Variable: Description: Selection: No	Drop list Vsc vs E plotve Allows to plot the correlation between spacecraft potential and electric field. No plot
Yes	Plots the correlation between spacecraft potential and electric field
Type: Label: Variable: Description: Selection: No Yes	Drop list Time series plottime Allows to plot time series of various quantities. No plot Plots time series of various quantities
Type: Label: Variable: Description: Selection: No Yes	Drop list Corr: E-I plotncorr Allows to plot the correlation between electron and ion density. No plot Plots the correlation between electron and ion density
Type: Label: Variable: Description: Selection: No Yes	Drop list nTlv donovertplot Allows to plot the dependence of current and density on various parameters. No such plots Plots current over temperature, current over electron velocity, current over ion velocity, current over ion Mach number, density over temperature, density over electron velocity, density over ion velocity, electron velocity over temperature, ion velocity over temperature.
Type: Label: Variable: Description: Selection: No Yes	Drop list Vsc-Tv dovscovervtplot Allows to plot the dependence of spacecraft potential on various parameters. No such plots Plots spacecraft potential over electron velocity, ion velocity, temperature, and ion Mach number.

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Туре:	Drop list
Label:	Detrend
Variable:	dodetrend
Description:	If set, a trend of log(current) with potential is subtracted from the current, and the result is set to 1 at the mean potential.
Selection:	
No	No detrend
Yes	A detrend as described above is done.
Type	Dron list

meas-rec-corr w/T
domeasreccorr
Defines whether the correlation between measured and reconstructed density
for a given temperature is plotted
No such plots
Measurement-Reconstruction correlation is plotted.

5.1.5 Line 5

In this line the limits of various solar activity parameters can be set, which are used for filtering the input data. In addition, the limits for the tilt angle of the spacecraft axis to the sun can be set.

Type:	Number
Label:	Solar activity: Irradiance>
Variable:	seemin
Description:	Sets the lower limit of SEE in [W/m^2/nm].
Type:	Number
Label:	<
Variable:	seemax
Description:	Sets the upper limit of SEE in [W/m^2/nm].
Type:	Number
Label:	F10.7 cm Flux>
Variable:	f107min
Description:	Sets the lower limit of F10.7 cm flux in [Jy].
Type:	Number
Label:	<
Variable:	f107max
Description:	Sets the upper limit of F10.7 cm flux in [Jy].
Type:	Number
Label:	Sunspots>
Variable:	sunspotmin
Description:	Sets the lower limit of the sunspot number.

Туре:	Number
Label:	<
Variable:	sunspotmax
Description:	Sets the upper limit of the sunspot number.

Type: Label:	Number Tilt(deg)>
Variable:	tiltmin
Description:	Sets the lower limit of the tilt angle.

т.

Туре:	Number
Label:	<
Variable:	tiltmax
Description:	Sets the upper limit of the tilt angle.

5.1.6 Line 6

In this line several parameters related to orbit filtering and plotting are defined

Type: Label: Variable: Description:	Number Orbit: R> orbitrmin Sets the lower limit of the radial distance from the Earth in Earth radii, which is used for filtering the input data.
Type: Label: Variable: Description:	Number < orbitrmax Sets the upper limit of the radial distance from the Earth in Earth radii, which is used for filtering the input data.

Type: Label: Variable: Description:	Number Phase(Sun=0)> orbitphasemin Sets the lower limit of the angular location in the GSE XY plane in degrees.
	Zero points to the Sun, the angle is defined in anticlockwise direction. For example, to select data from within 10° form the Sun enter 350° in this field and 10° in the field to the right.
Туре:	Number

number
<
orbitphasemax
Sets the upper limit of the angular location in the GSE XY plane.

Type:	Drop list
Variable: Description: Selection:	anyinoutbound Selects the inbound, outbound, or both legs of the orbit.
In- & Outbo	All parts of the orbit
Inbound or	ly Select only the inbound part of the orbit
Outbound o	only Select only the outbound part of the orbit
Type: Label:	Drop list Orbit map
Description: Selection:	Selects the parameter for the colour scale in the orbit map.
None E-field-tota	No orbit map is plotted I
SC-Potenti	Parameter is the total electric field al
Density Temperatu	Parameter is the spacecraft potential Parameter is the density of the species selected above re
Current	Parameter is the temperature of the species selected above
ASPOC Cu	Parameter is the current of the species selected above irrent
Modified C	Parameter is the ASPOC current urrent Parameter is the current of the species selected above in a modified
EL Velocity	calculation method
Ion Velocity	Parameter is the electron velocity
Debye Len	, Parameter is the ion velocity gth
Ion Mach N	Parameter is the Debye length lumber
Rel Ve Erro	Parameter is the ion Mach number
	Parameter is the relative error of the electron velocity in the input data. Not available in the current version of the input files.
Rel Vi Erro	r Parameter is the relative error of the ion velocity in the input data. Not available in the current version of the input files.

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Type:	Drop list
Label:	Scale
Variable:	posscaleislimits
Description: Selection:	Selects the range of the colour scale in the orbit map.
Data	The range of the colour scale is defined by the input data, additionally limited to the range given in group 5.2.
Limits	The range of the colour scale is defined by the limits specified in the control panel

Туре:	Drop list
Label:	Resol.
Variable:	posbini
Description:	Selects the spatial resolution of the orbit map.
Selection:	
0.1Re	One pixel is 0.1 x 0.1 Earth radii
0.2Re	One pixel is 0.2 x 0.2 Earth radii
0.5Re	One pixel is 0.5 x 0.5 Earth radii
1Re	One pixel is 1 x 1 Earth radii
2Re	One pixel is 2 x 2 Earth radii
4Re	One pixel is 4 x 4 Earth radii
	_
Type:	Number
Label:	Central Msheath@X&+Y&-YGSE
Variable [.]	xasedist
Description:	Sets the X(GSE) value of the nominal parabola defining the central
	magnetosheath position in the GSE XY plane. for $Y(GSE)=0$.
	-
Turner	Number
Type:	Number
	nygoodiat
Variable.	pyyseusi Sate the positive V(CSE) value of the pominal parabola defining the control
Description.	magnetosheath position in the CSE XX plane for X(CSE)-0
	magnetosneath position in the GSE \times 1 plane, for \times (GSE)=0.
	-
Type:	Number
	nyasedist
Description:	Sets the pagative V(GSE) value of the nominal parabola defining the central
Description.	magnetosheath position in the GSE XY plane for X(GSE)=0
	-

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Type:	Drop list
Label:	Region
Variable:	regionselect
Description: Selection:	Selects the region in space for which the input data are filtered.
All	All regions are selected (no filtering)
Msph	Regions inside the magnetosphere are selected. For the definition of boundaries see label "Width@Y=0&Y=20".
Wind	Regions in solar wind are selected. For the definition of boundaries see label "Width@Y=0&Y=20".
Sheath	Regions inside the magnetosheath are selected. For the definition of boundaries see label "Width@Y=0&Y=20".

Type: Label: Variable:	Number Width@Y=0&Y=20 regiontrans0 This is the width of the transition range (1/) around the nominal control
	magnetosheath position at Y(GSE)=0.
Type: Label:	Number
varianie'	regioniranszu

Variable:regiontrans20Description:This is the width of the transition range (+/-) around the nominal central
magnetosheath position at Y(GSE)=20 Earth radii.

5.1.7 Line 7

Type:	Drop list Color label is
Variable [.]	colorfrom
Description:	Selects the parameter for the colour scale in the correlation plots
Selection:	
None	No colour scale is applied
E-field-tota	
	Parameter is the total electric field
SC-Potenti	al
	Parameter is the spacecraft potential
Density	Parameter is the density of the species selected above
Temperatu	re
	Parameter is the temperature of the species selected above
Current	
	Parameter is the current of the species selected above
ASPOC CL	irrent
	Parameter is the ASPOC current
Modified C	Urrent
	Parameter is the current of the species selected above in a modified
Padius	
Naulus	Parameter is the Farth distance
XY GSE A	nale
	Parameter is the angle in the GSE XY plane
Rel. Time	
	Parameter is the time
Solar Index	(
	Parameter is the previously selected solar activity index
El. Velocity	,
	Parameter is the electron velocity
Ion Velocity	
	Parameter is the ion velocity
Debye Len	gtn Developmenter in the Debug loopsth
lon Moch N	Parameter is the Debye length
ION MACH P	Parameter is the ion Mach number
Rol Va Erro	
	Parameter is the relative error of the electron velocity in the input data. Not
	available in the current version of the input files.
Rel Vi Erro	r
	Parameter is the relative error of the ion velocity in the input data. Not
	available in the current version of the input files.
Туре:	Drop list
Label:	Colour is occurrence
Variable:	doocc
Description:	Defines the colour of the individual bins in the alternative plots of density or
Soloction	current over spacecraft potential (extensions _IVSCD and _NVSCD).
No	The bins are coloured according to the selected colour labol above
Yee	The bins are coloured according to the occurrence of the data
	The sine are belowed according to the bootheries of the data.

Drop list Label time plots dotimecolor Selects the style of time series plots. Time series plot style is full lines in a single colour.
Time series are plotted with symbols in the colour according to the previously selected parameter.
Number #labels in n-V-plots ntimelabels Defines the maximum number of labels attached to data points in the density over potential plots. The maximum number is 100. Note that the parameter range within the labels is defined in the fields at the right.
Number in V
labelvscmin Defines the minimum spacecraft potential in the density-potential plots which is labelled.
Number
labelvscmax Defines the maximum spacecraft potential in the density-potential plots which is labelled.
Number
labeldensmin Defines the minimum particle density in the density-potential plots which is labelled.
Number
labeldensmax Defines the maximum particle density in the density-potential plots which is labelled.
Drop list Scale ranges dofixscales Selects the scale range style in all plots except the orbit plots.
The scale range is defined by the input data The scale range is defined by the limits specified in the control panel

Туре:	Drop list
Label:	2nd param f. solar correl.
Variable:	secondvar
Description:	Defines whether an additional dimension shown as coloured symbols is
	applied in correlations with solar activity.
Selection:	
None	No second variable is used
E-field-tota	
	Second variable is the total electric field
SC-Potenti	al
	Second variable is the spacecraft potential
Density	Second variable is the density of the species selected above
Temperatu	re
	Second variable is the temperature of the species selected above
Current	
Carronic	Second variable is the current of the species selected above
ASPOC CI	irrent
	Second variable is the ASPOC current
Modified C	
	Second variable is the current of the species selected above in a modified
	calculation method
Dediue	
Raulus	Casend veriable is the Forth distance
XY GSE AI	
	Second variable is the angle in the GSE XY plane

5.1.8 Line 8

Type:	Drop list
Label:	Correlate Vsc with
Variable:	vsccorrwith
Description:	Defines whether an additional dimension shown as coloured symbols is applied in correlations with solar activity.
Selection:	
None Density	No correlation of any parameter is performed with spacecraft potential Spacecraft potential is correlated with particle density
Density and	Spacecraft potential is correlated with particle density and current
Density and	Spacecraft potential is correlated with particle density, current, and a current calculated in a modified way
Temperatu	re
Electric fiel	Spacecraft potential is correlated with temperature of selected species d
	Spacecraft potential is correlated with the total electric field
Uncontr. V	sc at ASPOC ONOFF or from 2 SC
	If data of only a single spacecraft are available, then a correlation between controlled and uncontrolled potentials around times when ASPOC turned on or off is performed. If data of two spacecraft are available, then the potential of the primary spacecraft is assumed to be the controlled one and the other potential the uncontrolled one, and a correlation between controlled and uncontrolled potential is performed.

Type: Label: Variable: Description:	Drop list As j Use tempwithvel Defines whether the current in the input file is replaced by a combination of thermal current and current from bulk flow
Selection: j(th) j(th)+j(ve) j(th)+j(vi) SQRT(j(th) SQRT(j(th)	No change from the input values. The current used in the calculations is the thermal current of the input file. The current used in the calculations is the sum of the thermal current of the input file and the electron current of derived from the bulk flow. The current used in the calculations is the sum of the thermal current of the input file and the ion current of derived from the bulk flow. $^{2}+j(ve)^{2}$ The current used in the calculations is the geometric sum of the thermal current of the input file and the electron current of derived from the bulk flow. $^{2}+j(ve)^{2}$ The current used in the calculations is the geometric sum of the thermal current of the input file and the electron current of derived from the bulk flow. $^{2}+j(vi)^{2}$ The current used in the calculations is the geometric sum of the thermal current of the input file and the ion current of derived from the bulk flow.
Type: Label: Variable: Description: Selection: Total SpinPlane Axial X-GSE Y-GSE Z-GSE	Drop list As v Use velocomp Defines which components of the bulk velocity are used for the calculations. The total velocity is used. The total velocity components in the spin plane are used. The total velocity component in axial direction is used. The total velocity component in X(GSE) direction is used. The total velocity component in Y(GSE) direction is used. The total velocity component in Z(GSE) direction is used.
Type: Label: Variable: Description:	Number with v times tempwithvelfudge Sets an arbitrary factor by which the bulk velocity is multiplied to obtain the current.

Type: Label: Variable: Description:	Number ASPOC step analysis: time window(s) tawindow This entry specifies the time window used in the comparison between controlled and uncontrolled potentials around a turn-on or turn-off of ASPOC. The program will look for all pairs of data which are tawindow seconds apart. There will be several pairs for any single on/off. The program will determine the significant differences.

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Туре:	Number
Label:	I-window
Variable:	iawindow
Description:	This entry specifies the minimum change of ASPOC current (in μ A) that characterizes a turn-on or turn-off.

Type:	Number
Label:	Max gap
Variable:	maxgap
Description:	This entry specifies the maximum allowed size of a data gap (i nunits of spin periods) around an apparent turn-on or turn-off of ASPOC to be considered in the analysis.

5.1.9 Line 9

Type: Label:	Number n &I-T-exponent:
Variable:	novertexp2
Description:	This is the exponent to temperature used to calculate plasma current under certain conditions, see below.

Type: Nu	umber
Label: ap	oplies for T<
Variable: no	overttemp2
Description: Th	nis is temperature T2 (in eV) used to calculate plasma current under certain
co	onditions, see above.

Type:	Number
Label:	and exp:
Variable:	novertexp
Description:	This is the exponent to temperature used to calculate plasma current under certain conditions, see above.

Type: Label: Variable: Description:	Number for higher T up to: noverttemp The particle current is calculated as being proportional to density times the square root of temperature for temperatures larger than set in this field. If a value greater than 0 is specified, then the standard formula for calculating current (proportional to density times the square root of temperature) is not applied for a certain range of temperatures. This field and the two fields at the right (labeled "T-exponent" and "for T<") work together. Temperatures are given in eV. Let us call this temperature T1 and the one at right T2.
	Case T1>0 AND T1 <t2: Measured T<t2: and="" by="" calculated="" current="" exponent="" fixed="" is="" t1.<br="" the="">Measured T>T2: Current is taken from the input file where the standard formula has been used</t2:></t2:
	Case T1>0 AND T1>T2: Measured T <t2: and="" by="" calculated="" current="" exponent="" fixed="" is="" t1.<br="" the="">T1>Measured T>T2: Current calculated by the standard formula. Measured T>T1: Current is taken from the input file where the standard formula has been used.</t2:>
	Case T1>0 AND T1=T2: Measured T <t2: and="" by="" calculated="" current="" exponent="" fixed="" is="" t1.<br="" the="">Measured T>T2: Current is taken from the input file where the standard formula has been used.</t2:>
Type: Label: Variable: Description:	Number Enhance n&I at T< plasmaenhanceT0 Densities and currents below this temperature are modified by the factor plasmaenhanceT0^plasmaenhanceexp.
Type: Label: Variable: Description:	Number by plasmaenhanceexp Densities and currents below this temperature are modified by the factor plasmaenhanceT0^plasmaenhanceexp.
Type: Label: Variable: Description:	Number Min T for I-calc tempminforl This defines a minimum temperature used to calculate reconstructed densities.

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Drop list FPI spectra dospectra Selects whether an energy-time spectrogram of electron or ion flux is plotted. No plot Spectrogram is plotted
Number Max energy step maxestep Sets the maximum energy step of the FPI data which is plotted in the spectrogram. The input data contain 32 steps. This selection is relevant only if a partial moments file has been read.
Drop list Regres w/V dopartcorr Selects whether spacecraft potential shall also be correlated with partial particle density. This selection is relevant only if a partial moments file has been read. No correlation Correlation is performed

5.1.10 Line 10

Туре:	Drop list
Label:	Manual correction of ivsc
Variable:	doivsccorr
Description:	Selects the fitting function between current and spacecraft potential shall be corrected manually.
Selection:	
No	No manual correction
Yes	Apply manual correction

Type: Label: Variable: Description:	Number # points ivsccorrn Specifies the number of valid data points in the polynomial for the manual correction of the current-spacecraft potential fit. Minimum number: 2, maximum number: 6
Type: Label: Variable: Description:	Number Vsc ivsccorrVsc Specifies up to 6 spacecraft potential values in the polynomial for the manual correction of the current-spacecraft potential fit.

Type: Label:	Number Factors
Description:	Specifies up to 6 correction factors in the polynomial for the manual correction of the current-spacecraft potential fit. Corrected current = Automatically fitted current / correction factor.
5.1.11 Li	ne 11
Туре:	Drop list
Label:	Histograms of
Description: Selection:	Defines the parameter used to plot histograms versus a second parameter.
None E-field-tota	No histogram of any parameter is plotted
	Histograms of total electric field versus a second parameter are plotted
SC-Potenti	
Density Temperatu	Histograms of spacecraft potential versus a second parameter are plotted Histograms of particle density versus a second parameter are plotted
remperato	Histograms of temperature versus a second parameter are plotted
Current ASPOC Cu	Histograms of plasma current versus a second parameter are plotted
Modified C	Histograms of ASPOC current versus a second parameter are plotted urrent
Radius	Histograms of modified plasma current versus a second parameter are plotted Histograms of radial distance versus a second parameter are plotted
XY GSE A	ngle
	Histograms of the position angle in the GSE XY plane versus a second parameter are plotted
Solar Index	
	An are plotted
EI. Velocity	/ Histograms of electron bulk velocity versus a second parameter are plotted
Ion Velocit	y Histograms of ion bulk velocity versus a second parameter are plotted
Debye Len	gth
Ion Mach N	Histograms of Debye length versus a second parameter are plotted
	Histograms of ion Mach number versus a second parameter are plotted

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Type: L abel:	Drop list vs
Variable:	vschistsel
Description: Selection:	Defines the second parameter in the histograms.
None E-field-tota	No second parameter
	Second parameter is total electric field
SC-Potent	ial
	Second parameter is spacecraft potential
Density	Second parameter is particle density
Temperatu	Ire
	Second parameter is temperature
Current	Second parameter is plasma current
ASPOC C	urrent
	Second parameter is ASPOC current
Modified C	Current
	Second parameter is modified plasma current
Radius	Second parameter is radial distance
XY GSE A	ngle
	Second parameter is the position angle in the GSE XY plane
Solar Inde	X
	Second parameter is the solar index selected in line 4
EI. Velocity	
	Second parameter is electron bulk velocity
Ion Velocit	
Debuglar	Second parameter is ion bulk velocity
Debye Ler	Igin Second percenter in Debug length
lan Maah I	Second parameter is Debye length
ION Mach	Number Second parameter is ion Mach number
-	-
Type:	Number

Number
#bins
histnbins
Specifies the number of bins in the histograms.

Туре:	Number
Label:	#param steps
Variable:	histnvstep
Description:	Specifies the number of curves in the histograms.

Type:	Number
Label:	Scatterplots 1:x points
Variable:	n100
Description:	Scatterplots may become too crowded if all data of a long time period are plotted. In this field the user may specify that only a subset of points is plotted.

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Label: Variable: Description:	Number Symbol size symsi Specifies the symbol size in scatterplots.
Type: Label: Variable: Description:	Number #bins nscatbins Specifies the number of bins in alternative scatterplots, for example for der over spacecraft potential.
Type: Label: Variable: Description:	Number cutoff lowcutoff Specifies the minimum number of data points in each plotted bin in alterna scatterplots, for example for density over spacecraft potential. Bins with less data points are set emptynumber of bins in alternative scatterplots, for example for density over spacecraft potential.
Type: Label: Variable: Description:	Number #equal-weight bins nhistbins Specifies the number of bins of equal weight. This entry is valid only if in th drop list "Error exp" the value "hist" has been selected.
Type: Label: Variable: Description: Selection: No Yes	Drop list Omit data near NaN omitnearnan Selects whether data points in the time series which are adjacent to not available data (NaN) are omitted No omission Data are omitted

Type:	Number
Label:	Corrections for Vsc: E-field[km]
Variable:	efldfac
Description:	This value defines the trend correction factor of spacecraft potential for the electric field, $Vsctrend[V] = E[mV/m]^*$ factor.

Туре:	Number
Label:	vi[mV s/m]
Variable:	vifac
Description:	This value defines the trend correction factor of spacecraft potential for ion bulk velocity in units of $[mV \ s/m]$.

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Type: Label: Variable: Description:	Number Sun sunfac This value defines the trend correction factor of spacecraft potential for the solar activity parameter selected in line 4.
Type: Label: Variable: Description:	Number Mach vmachfac This value defines the trend correction factor of spacecraft potential for the ion Mach number.
Type: Label: Variable: Description:	Number Ig(T)Io tempfaclo This value defines the trend correction factor of spacecraft potential for the logarithm of temperature below the threshold temperature at the right.
Type: Label: Variable: Description:	Number Ig(T)hi tempfachi This value defines the trend correction factor of spacecraft potential for the logarithm of temperature above the threshold temperature at the right.
Type: Label: Variable: Description:	Number Tthresh tempfacthresh This value defines the threshold temperature for the two correction factors defined at the left.
Type: Label: Variable: Description:	Number Fudge factor for maxcurrent maxcurrentfactor The program uses a constant ASPOC current for some calculations, which is derived from the mean value of this current in the data. This field contains a correction factor to this.
Type: Label: Variable: Description:	Number for Iplasma iefudgefactor This field contains a correction factor for the plasma current calculated from density and temperature.

5.1.13 Line 13

Type: Label: Variable:	Drop list Show VV power fit in IV-plot doplotwresult
Description:	Defines the photocurve with power law parameters derived from the comparison of controlled and uncontrolled potential in a previous session shall be plotted in the present plot.
Selection: No Power	The previous photocurve is not plotted The photocurve is shown as power law, as previously calculated
Single Max	The previously calculated power law photocurve is approximated by a single Maxwellian function and plotted
Sum of 2 N	Aaxw The previously calculated power law photocurve is approximated by a sum of 2 Maxwellian functions and plotted
Sum of 3 N	Iaxw The previously calculated power law photocurve is approximated by a sum of 3 Maxwellian functions and plotted
Type: Label: Variable: Description:	Number coefs jph aph0 Factor in the power law approximation of photoelectron current density.
Type: Label: Variable: Description:	Number bph1 bph1 Exponent in the power law approximation of photoelectron current density.
Type: Label: Variable: Description:	Number Use Vsc dev. from smoothing x spins (0=no). x: smoothvscdata If set greater than 0, this value sets the smoothing range of the spacecraft potential in the correlation with the electric field.
Type: Label: Variable: Description:	Number Smooth dV/dE (0=no). x: smoothdvdedata If set greater than 0, this value sets the smoothing range of dVsc/dE in the plot of this quantity over time.
Type: Label: Variable: Description:	Number dV/dE scale< dvdemax This entry sets the scale maximum of the quantity dVsc/dE in the plot of this quantity over time.

Type:	Drop list
Label:	Divide dV/dE by V
Variable:	dorelativedvde
Description:	Defines whether the dependence of spacecraft potential with the electric field is plotted in a normalized way, i.e. divided by the spacecraft potential.
Selection:	
No	Quantity dV/dE is plotted as is
Yes	Quantity is divided by spacecraft potential

Туре:	Drop list
Label:	Output photocurve
Variable:	dophotocurve
Description:	Defines whether the derived Current-Voltage curve (also named photocurve) is output as a plot and as text in the IDL status window.
Selection:	
No	No output
Yes	Output of photocurve

5.1.14 Line 14

Type: Label:	Drop list Fit I or n vs V on set polynomial not data
variable:	nipolynoidala
Description:	Defines whether the fitting in the current or density over spacecraft potential data shall be performed on the data or on the polynomial defined by the coefficients at the right.
Selection:	
No Yes	Fits are performed on data Fits are performed on the polynomial defined at the right

Туре:	Number
Label:	#coefs
Variable:	npolycoefs
Description:	Defines the number of polynomial coefficients approximating the n-V or I-V curves. Only the given number of coefficients at the right are applied.

Type:	Number
Label:	coefs:
Variable:	polycoef0
Description:	Coefficient 0 of the polynomial approximating the measured data.
Type:	Number
Variable:	polycoef1
Description:	Coefficient 1 of the polynomial approximating the measured data.

Type: Label:	Number
Variable: Description:	polycoef2 Coefficient 2 of the polynomial approximating the measured data.
Type:	Number
Variable: Description:	polycoef3 Coefficient 3 of the polynomial approximating the measured data.
Type:	Number
Variable: Description:	polycoef4 Coefficient 4 of the polynomial approximating the measured data.
Type:	Number
Variable: Description:	polycoef5 Coefficient 5 of the polynomial approximating the measured data.
Type:	Number
Variable: Description:	polycoef6 Coefficient 6 of the polynomial approximating the measured data.
Type:	Drop list
Variable:	doplotpoly Defines whether the polynomial fited to the data is plotted with the data
Selection:	Polynomial is not plotted together with the data
Yes	Polynomial is plotted together with the data
Type:	Drop list Use set low-E Maxw for Maxw-fit
Variable:	addmaxw
Description:	Defines whether the terms for the Maxwellian population at low energy given at the right shall be added when fitting the photocurve.
Selection: No	Pre-set low energy Maxwellian terms are not added
Yes	Pre-set low energy Maxwellian terms are added
Type: Label:	Number coefs: i0
Variable:	addj0
Description:	Factor of the low energy Maxwellian term added to the fit.

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Type: Label:	Number V0
Variable: Description:	addv0 Characteristic energy of the low energy Maxwellian term added to the fit.
5.1.15 Li	ne 15
Type: Label:	Drop list Fit I vs V
Variable: Description:	powerrmsi Defines the function used to fit the current over spacecraft potential data to obtain the photocurve.
Selection: No	No fitting is performed
Maxwellian	S
	A single Maxwellian distribution or a sum of 2 or 3 Maxwellian distributions is used for the fitting. The number of terms is given by the entry "#Terms" at the right.
Power i~V⁄	`x
Max w pow	A power law distribution i~V^x is used for the fitting. /er init
	A single Maxwellian distribution or a sum of 2 or 3 Maxwellian distributions is used for the fitting, using a power law fit for the start value of the nonlinear Maxwellian fit. The number of terms is given by the entry "#Terms" at the right.
Max fm pov	wer
	A single Maxwellian distribution or a sum of 2 or 3 Maxwellian distributions is used for the fitting. The fit is not performed on the data, but rather on the approximation of the power law fit. The number of terms is given by the entry "#Terms" at the right.
3-range Ma	ах
	Three Maxwellian terms separated in spacecraft potential are used for the fitting.
Power V~i/	^x
Polynomial	A power law distribution V~i/x is used for the fitting.
Tolynomia	A polynomial in current i is with the number of terms given at the right used for the fitting.
Type:	Drop list Maxw fit
Variable [.]	dofullmaxw
Description:	Defines the function used to fit the current over spacecraft potential data to obtain the photocurve.
Coloction	

Selection:

- The Maxwellian fit is performed without any weighting or limits by a simple regression function. The Maxwellian fit is performed by the IDL routing MPFIT including weights Regress
- Mpfit and limits.

Type:	Number
Label:	#Terms
Variable:	ntermsrms
Description:	Number of terms used in the fitting of Maxwellian functions.
Type: Label: Variable: Description: Selection: No Yes	Drop list Limits limitedrms Defines whether the non-linear fitting of the Maxwellian terms shall be performed with limits to the parameters. non-linear fitting of Maxwellian functions are performed without limits for parameters non-linear fitting of Maxwellian functions are performed out limits for parameters
Type:	Drop list
Label:	Method
Variable:	methodrms
Description:	Defines the formula used to calculate plasma electron current. The "exact"
Selection:	option includes the correction terms for an attracting sphere
Simple	Current ~ density * root(temperature)
Exact	Current ~ density * root(temperature) * (1 + potential/temperature_in_eV)
Type: Label: Variable: Description: Selection: -3 -2 -1 0 0.5 1 2 hist	Drop list Error exp weightsrmse The convergence of the nonlinear fit to obtain the photo-emission spectrum is critical. This selection allows to define relative weights to the data points to improve the convergence. Weight is uncontrolled potential (of the second spacecraft) to the power -3. Weight is uncontrolled potential (of the second spacecraft) to the power -2. Weight is uncontrolled potential (of the second spacecraft) to the power -1. Weight is uncontrolled potential (of the second spacecraft) to the power -1. Weight is uncontrolled potential (of the second spacecraft) to the power -0.5. Weight is uncontrolled potential (of the second spacecraft) to the power +1. Weight is uncontrolled potential (of the second spacecraft) to the power +2. The total weight of all data points of the uncontrolled potential (of the second spacecraft) in each interval is equal. There are 30 intervals, logarithmically spaced in the range between 1 V and 60 V.

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Type: Label:	Drop list in	
Variable: Description:	weightsinye Defines whether the error is calculated in the Y axis (dependent variable) or X avia (independent variable)	
Selection:		
X Y	Errors are calcualted for the independent variable, usually plotted in X Errors are calcualted for the dependent variable, usually plotted in Y	
Type: Label: Variable: Description:	Drop list Break V Vbreakvariab Defines whether the potentials separating the validity of three Maxwellian fittings are set manually by the values entered at the right, or are calculated	
Coloction	automatically based on minimum total error.	
Fix Var	Fixed limits given at the right are used Variable limits are calculated	
Type: Label:	Number	
Variable: Description:	Vbreak1e Potential separating the fits at low and medium potential	
Type:	Number	
Variable: Description:	Vbreak2e Potential separating the fits at medium and high potential	
Type: Label: Variable: Description:	Drop list T for n-fit varassumedtemp Defines whether the temperature assumed for the conversion of density to current is assumed to be constant at the value given at the right under the label "fix", or a function of spacecraft potential given by a straight line between	
Selection:	two points given at the right.	
Fix Var	A fixed temperature is assumed A variable temperature is assumed	
Type: Label: Variable: Description:	Number at 2 V assumedt0 Assumed variable temperature at 2 V spacecraft potential	

Type:	Number
Label:	7 V
Variable:	assumedt1
Description:	Assumed variable temperature at 7 V spacecraft potential
Type:	Number
Label:	10.5 V
Variable:	assumedt2
Description:	Assumed variable temperature at 10.5 V spacecraft potential
Type:	Number
Label:	20 V
Variable:	assumedt3
Description:	Assumed variable temperature at 20 V spacecraft potential
Туре:	Number

1,900.	
Label:	fix:
Variable:	assumedtemp
Description:	Assumed constants temperature

5.2 Group "SETTINGS FOR SPACECRAFT 1"

In this part of the screen the settings, mainly filters, applied to the primary spacecraft - or in general if no secondary spacecraft has been selected - are defined.

5.2.1 Line 1	5.2.1	Line	1
--------------	-------	------	---

Type:	Number
Label:	Efield>
Variable:	efldmin
Description:	Minimum total electric field used for filtering input data and for setting the scale range.

Туре:	Number
Label:	>
Variable:	efldmax
Description:	Maximum total electric field used for filtering input data and for setting the scale range.

Type:	Number
Label:	Vsc>
Variable:	vscmin
Description:	Minimum spacecraft potential used for filtering input data and for setting the scale range.

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Type: Label: Variable: Description:	Number > vscmax Maximum spacecraft potential used for filtering input data and for setting the scale range.
Type: Label: Variable: Description: Selection: Linear Logarithmic	Drop list Vsc scale dologvsc Defines the scale for the spacecraft potential in the correlation plots. Spacecraft potential is plotted in a linear scale.
Type: Label: Variable: Description:	Number laspoc> aspmin Minimum ASPOC current used for filtering input data. Note that the data files in the distribution contain the difference between the ASPOC current and the EDI current. Therefore, in oder to include all data with ASPOC OFF, the minimum current should be set to -1 and the maximum current to +1.
Type: Label: Variable: Description:	Number < aspmax Maximum ASPOC current used for filtering input data. Note that the data files in the distribution contain the difference between the ASPOC current and the EDI current. The nominal ASPOC current is 20µA per spacecraft. Therefore, in oder to include all data with ASPOC ON excluding the current sweeps, the minimum current should be set to 19 and the maximum current to +21.
Type: Label: Variable: Description:	Number Density: ne-ni < densdiffmax Sets the maximum permitted difference between measured electron and ion density.
Type: Label: Variable: Description:	Number AND ne-ni < densfactmax Sets the maximum permitted factor between measured electron and ion density.
Type: Label: Variable: Description:	Number Te for fit> etempminforfit Sets the minimum temperature used for fitting current or density.

Type: Label: Variable: Description:	Number IonMach> machmin Sets the minimum ion Mach number used for filtering input data.
Type: Label: Variable: Description:	Number > machmax Sets the maximum ion Mach number used for filtering input data.
Type: Label: Variable: Description:	Number Min v-comp velocompmin sets the minimum value of individual bulk velocity components used for filtering input data
5.2.2 Li	ne 2
Type: Label: Variable: Description:	Number Electrons: Ne> edensmin Minimum electron density used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number > edensmax Maximum electron density used for filtering input data and for setting the scale range.
	-

Type:	Number
Label:	Te>
Variable:	etempmin
Description:	Minimum electron temperature used for filtering input data and for setting the scale range.

Type:	Number
Label:	>
Variable:	etempmax
Description:	Maximum electron temperature used for filtering input data and for setting the scale range.

Type:	Number
Label:	le>
Variable:	ecurrmin
Description:	Minimum electron current used for filtering input data and for setting the scale
	range.

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Type: Label: Variable: Description:	Number > ecurrmax Maximum electron current used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number ve> evelomin Minimum electron velocity used for filtering input data.
Type: Label: Variable: Description:	Number > evelomax Maximum electron velocity used for filtering input data.
Type: Label: Variable: Description:	Number Ions: Ni> idensmin Minimum ion density used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number > idensmax Maximum ion density used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number Ti> itempmin Minimum ion temperature used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number > itempmax Maximum ion temperature used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number li> icurrmin Minimum ion current used for filtering input data and for setting the scale range.

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Type: Label: Variable: Description:	Number > icurrmax Maximum ion current used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number vi> ivelomin Minimum ion velocity used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number < ivelomax Maximum ion velocity used for filtering input data and for setting the scale range.

5.3 Group "SETTINGS FOR SPACECRAFT 2 OR FOR UNCONTROLLED DATA IF APPLICABLE"

In this part of the screen the settings applied to the secondary spacecraft - if present - are defined.

5.3.1 Line 1

Type:	Drop list
Label:	Plot Vsc2(unc) over Vsc1(cont) OR Vsc1 at ASPOC changes
Variable:	dovvcorr
Description:	Defines whether a correlation between controlled and uncontrolled spacecraft potential shall be performed.
Selection:	
No	No correlation
Yes	Correlation is performed

Type:NumberLabel:Iaspoc2>Variable:asp2minDescription:Minimum ASPOC current in the uncontrolled case for either the primary or
secondary spacecraft used for filtering input data. Note that the data files in
the distribution contain the difference between the ASPOC current and the
EDI current. Therefore, in oder to include all data with ASPOC OFF, the
minimum current should be set to -1 and the maximum current to +1.

Type: Label: Variable: Description:	Number < asp2max Maximum ASPOC current in the uncontrolled case for either the primary or secondary spacecraft used for filtering input data. Note that the data files in the distribution contain the difference between the ASPOC current and the EDI current. Therefore, in oder to include all data with ASPOC OFF, the minimum current should be set to -1 and the maximum current to +1.
Type: Label: Variable: Description:	Number (Vsc2 only) Vsc2> vsc2min Minimum spacecraft potential current in the uncontrolled case for either the primary or secondary spacecraft used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Number > vsc2max Maximum spacecraft potential current in the uncontrolled case for either the primary or secondary spacecraft used for filtering input data and for setting the scale range.
Type: Label: Variable: Description:	Drop list Vsc2 scale dologvsc2 Defines the scale for the uncontrolled spacecraft potential in the correlation plots.
Linear Linear Logarithmi	Uncontrolled spacecraft potential is plotted in a linear scale. c Uncontrolled spacecraft potential is plotted in a logarithmic scale.
Type: Label: Variable: Description:	Number Vsc2 correction vsc2correct Arbitrary offset added to the potential of the secondary spacecraft.

5.3.2 Line 2

Type: Label: Variable: Description: Selection: None 2-range Ma	Drop list Fit dompfitv Defines the function used to derive the photocurve form controlled and uncontrolled potential. No fitting is performed axw Two Maxwellian terms separated in spacecraft potential are used for the fitting
Power	A power law distribution is used for the fitting.
Type: Label: Variable: Description:	Drop list Maxw break V Vbreakvariabv Defines whether the potential separating the validity of two Maxwellian fittings is set manually by the value entered at the right, or is calculated automatically based on minimum total error
Selection: Fix Var	The fixed limit given at the right is used A variable limit is calculated
Type: Label: Variable: Description:	Number Vbreak1v Potential separating the fits at low and high potential
Type: Label: Variable: Description:	Drop list Error exp weightsrmsv The convergence of the nonlinear fit to obtain the photo-emission spectrum is critical. This selection allows to define relative weights to the data points to improve the convergence.
Selection: -3 -2 -1 0 0.5 1 2 hist	Weight is uncontrolled potential (of the second spacecraft) to the power -3. Weight is uncontrolled potential (of the second spacecraft) to the power -2. Weight is uncontrolled potential (of the second spacecraft) to the power -1. Weight is unity. Weight is uncontrolled potential (of the second spacecraft) to the power -0.5. Weight is uncontrolled potential (of the second spacecraft) to the power +1. Weight is uncontrolled potential (of the second spacecraft) to the power +2. The total weight of all data points of the uncontrolled potential (of the second spacecraft) in each interval is equal. There are 30 intervals, logarithmically spaced in the range between 1 V and 60 V.

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Type: Label: Variable: Description:	Drop list in weightsinyv Defines whether the error is calculated in the Y axis (dependent variable) or X axis (indeendent variable).
Selection:	
X Y	Errors are calcualted for the independent variable, usually plotted in X Errors are calcualted for the dependent variable, usually plotted in Y
 Type:	Number
Label:	Start coefs
Variable: Description:	startcoefs0 Start value in the fit for the factor of the power law or of the first Maxwellian term, in μ A/m ⁻²
Туре:	Number
Label:	startcoofs1
Description:	Negative exponent of the power law or characteristic potential of the first Maxwellian term in V
Type: Label:	Number

Variable: startcoefs2

Description: Characteristic potential of the second Maxwellian term in V. This entry is not applicable for the power law fit.

6 References

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7 Annex: List of Routines

7.1 Main program

CORRVANDFLUX

7.2 Routines inside corrVandFlux.pro

EP_TICKSD GAPOPLOT GETIASPOC1D GETLNIASPOC3PARTS GETWC2PARTS GETVC2PARTS GETVC3PARTS GETVCPOWER NOTICK PRINTF111 POWERTOMAXW TIMEAXISD TWOSLOPES YGAPOPLOT

7.3 Third party routines

MPFITFUN